Astrophysics and Nuclear Astrophysics (LPHY2263)

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Chapter #9

- Introduction to Cosmology
 - Galactic red shifts and distances
 - Hubble Law
 - Expanding Universe
 - Big Bang
 - Quark-Gluon Plasma
 - Inflation
 - Accelerating Universe









Red shift / blue shift

Definition of *red shift / blue shift*:

$$z = \frac{\lambda_0 - \lambda_e}{\lambda_e} = \frac{\lambda_0}{\lambda_e} - 1$$

It has three components:

$$z_{\text{total}} = z_{\text{grav}} + z_{\text{physical}} + z_{\text{H}}$$

The first term (z_{grav}) is due to the effect from gravity: $\Delta \nu - GM (1 \ 1)$

$$\frac{\Delta v}{v} = \frac{-GM}{c^2} \left(\frac{1}{r_1} - \frac{1}{r_2}\right)$$

The second term (z_{phys}) is due to motion within galaxy / cluster / supercluster; the third term (z_{H}) is cosmological and we will try to understand it today

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First surprising discovery

- Between 1914 and 1922, Vesto Slipher used this technique to measure the speed of 25 galaxies
- He found that 21 out of 25 had a redshift, i.e., they go away from us
- Some of them with speeds of > 1000 km/s
- Local or global phenomenon? Not known, because distance was not measurable yet
- Hubble was able in 1929 to measure their distances, and notice a speed-distance relationship
- Next slides are about how we measure galactic distances over very large scales

Measurements of distances

- Parallax method cannot work for galaxies
 - (Do you remember why?)
- Methods that work are based on "standard candles", i.e., objects of known luminosity, for example:
 - Cepheids
 - Most luminous supergiants
 - Supernovae
 - Tully-Fisher relation



Picture from here

Cepheids



Cepheid variables can measured out to ~20 Mly (million light years)

Most luminous supergiant

- The brightest supergiants in a given galaxy will have about the same absolute magnitudes
 - -8 for red supergiants and -9 for blue supergiants
 - Visible from about 50 Mly and 80 Mly respectively
 - Method works because there is a well-defined maximum in path in the H-R diagram of a star that goes out of the Main Sequence



Remember chapters 4 and 5 of these lessons

Supernovae

- Beyond 300 Mly, good standard candles are the supernovae, which can reach a peak magnitude of -19
- In principle, such a magnitude should be visible to 8 billion light years
- Supernovae of types Ib, Ic and II are not good candles, because they have too much variety in L
- Type-Ia supernovae instead are very uniform, for the reason explained in the next slide

Type-la supernovae



Picture from here

Accretion is gradual, and explosion happens as soon as the mass passes the Chandrasekhar threshold of 1.44 M_{sun}. Important point: it always starts at the same mass. Therefore, the energy output is almost the same

Tully-Fisher relation



Empirical relation between a galaxy's intrinsic luminosity and the width of its emission lines, which is related to its rotation speed

The "distance ladder"



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Each step is validated and calibrated by the previous one and helps to calibrate the next one



- Deviations from the straight line are attributed to the fact that all galaxies have additional motion in addition to pure expansion
- This is called the "cosmic velocity dispersion" or "cosmic scatter"



- distance
- Here in the diagram there are a bunch of galaxies at around the same distance but spread out a lot in velocity
- Attributed to galaxies from the Virgo cluster: extra velocity from cluster orbital motion, that can be very large

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Type-Ia supernovae give very precise measurements:



Figure 1. The redshift–distance relation for type Ia supernovae from the Calan-Tololo supernova survey (from data courtesy of S Jha, Harvard-Smithsonian). Redshift is given as *cz* and the distance is in megaparsecs (1 Mpc is approximately 3.26 million light years). The line through the data represents a Hubble constant of 65 km s⁻¹ Mpc⁻¹, close to the value astronomers measure today.

Cepheids also help a lot at small distances; combination of ¹⁵ both methods gives ±5-10% precision on H

Expanding Universe



Picture from here

Questions

- This is an expansion of space itself; does it mean that galaxies become larger too?
 - (...and solar systems, and planets, and ourselves...)
- At sufficiently large distances, the velocity calculated from the redshift is larger than c; does it mean that relativity is wrong?
- If you assume that the v=Hd relationship was always true, how old is the Universe?

Big Bang

- Simply by extrapolating to the past, we are forced to think that at some point (at time 1/H) the current Visible Universe was very small, and therefore very dense
- George Lemaître (UCL) was the first to propose that all Universe started in a "singularity"
 - Einstein and others studied the consequences in General Relativity, based on two principles: universality of the physics laws, and the "cosmological principle", which states that the Universe is <u>homogeneous and isotropic at large scales</u>
- Not obvious that the singularity must be taken literally: it may be that a coherent theory of quantum gravity, when we will know it, will get rid of singularities
 - (e.g., it is plausible that the limit is a volume of size $\sim h^3$) 18

Alternative: the Steady State theory

- Not everybody was happy with a model that violates the intuitive idea that the Universe is invariant in time
- Fred Hoyle proposed an alternative, based on the principle that the Universe looks the same at any point in time ("perfect cosmological principle") → no start!
- To concile it with expansion, it meant that matter is continuously and uniformly created from vacuum
- The model started to loose popularity when we started to know phenomena (like quasars) that happened in the distant past but not anymore
- Then, the model was definitely dead when the following "smoking guns" of Big Bang were found

"Smoking guns" of the Big Bang

- The most striking prediction of the Big Bang concept is that, at the beginning, the Universe was very dense and hot
- The next slides present:
 - An indirect proof: primordial nucleosynthesis (it was hot enough to allow nuclear fusion)
 - A direct proof: Black Body radiation of the Universe

Primordial Nucleosynthesis

- There are indications that only part of the Helium in the Universe was produced by stellar nucleosynthesis:
 - Population I stars (e.g., the Sun): 70% H, 28% He, ~2% metals
 - Population II stars (i.e., poor in metals): 75% H, 25% He, < 0.01% metals (remember: "metal" means heavier than He, in astro jargon)
- Where does the He in Pop.II stars come from?
 - If it comes from previous generation stars, where are the metals?
- When the Universe had less than 2 minutes: too hot for nuclei
- Around ~2 minutes, T~O(10⁹ K) and started to be cold enough to allow ²H formation without immediate fission; all n were used
- Between 2 and 4 minutes, most ²H fused into ⁴He
- After 4 minutes, too cold for nuclear fusion (\Rightarrow no metals)

Cosmic Microwave Background (1)

- After nucleosynthesis, for 300k years nothing dramatic happens; in that period, Universe cooled down to 3000 K
- Before reaching that temperature, Universe was too hot to allow atoms to form
- Universe was a plasma of free electrons and nuclei (p, α)
 - Free electrons easily scatter photons
 - \Rightarrow Universe was completely opaque to light
 - \Rightarrow Perfect Black Body
- When T=3000 K, Universe suddenly became transparent
 - The oldest photons that we can observe come from that precise point in time, and were produced with a Black Body spectrum for T=3000 K; and then redshifted...

Cosmic Microwave Background (2)

- In 1965, Penzias and Wilson (telecommunication engineers at Bell Labs) mapped the microwave sky
- There was a faint isotropic noise
- They made very deep studies to understand the source of that noise
 - It was not electronic noise from the amplifiers
 - And not the pigeons that were "dirtying" the antennas
 - ...etc.
- In the end, they discovered that it was of cosmic origin
- It corresponds to a Black Body of 2.7 K



Quark-Gluon Plasma (QGP)

- There must have been a period in the history of the Universe (order of milliseconds) when temperature was so high that even protons and neutrons could not form
 - i.e., Universe was filled with a plasma of quarks and gluons
- The properties of QGP (e.g., T of phase transition; liquid or gas; etc.) affected the evolution of the Universe
- The problem is that we have no idea of the temperature of this phase transition
 - Astronomical observations can't tell us much, because this was much before the "first light"
 - Quantum Chromo-Dynamics (QCD) is not in the perturbative regime in these density conditions \rightarrow calculations are very imprecise, and based on semi-empirical models ²⁵

Large Hadron Collider (LHC)



ALICE experiment (A Large Ion Collider Experiment)



- One month per year, LHC accelerates heavy-ions (so far only Pb) instead of the usual protons
- The goal is to create, for a short time, a quark-gluon plasma and therefore study the properties of matter in a very young Universe
- The ALICE detector is optimized for those studies (although ATLAS and CMS are somewhat complementary also on that)

Pb-Pb collision in ALICE



Three main historical problems of the Big Bang theory

- Horizon problem
 - Although space may expand faster than c, Relativity imposes that no information can travel faster than c
 - If so, how comes that the Black Body temperature is homogenous • on so large scales? "Thermalization" requires exchange of energy
- Flatness problem
 - The geometry of the Universe seems very flat; but any departure from flatness should have been amplified
 - There is probably a global curvature, but several orders of magnitude less than predicted by General Relativity
- Magnetic monopoles
 - Grand Unification Theories predict magnetic monopoles; Universe must have enough temperature to create them 29

Inflation (1)

- An unknown scalar field permeates the Universe (like the Brout-Englert-Higgs field, but with a larger mass)
- At a certain temperature it underwent a phase transition, which released a lot of energy (analogy: think about ice↔water transitions)
 - This phase transition happened very early in the history of the Universe because the mass of the field's quanta is very large and therefore T is very large
- This sudden release of energy caused, for a short period, an exponential expansion of the Universe



Inflation (2)

- Horizon problem
 - Black Body temperature is homogenous on large scales just because those corresponded to a very small volume, such that it thermalized quickly
- Flatness problem
 - Departures from flatness were indeed amplified, but then space was so much inflated that global curvature became small again
- Magnetic monopoles
 - Grand Unification Theories created them in large quantities, indeed, but this huge expansion has diluted their density so much that they became extremely rare
 - (However, some experiments are still looking for them...)

Gravity waves from Inflation

- One of the prediction of Inflation is that a lot of energy was released in the form of gravity waves
- In General Relativity, gravity is a tensorial field (gravitons have spin=2)
- Waves of vectorial fields (spin=1) cause movements up and down, waves of tensorial fields cause torsions
- Recent excitation when the BICEP2 collaboration seemed to have measured this torsion of space:



• Status: not confirmed (foreground subtraction was wrong)

Accelerating expansion

- The Hubble constant is constant only in space, not in time
- The redshift data on the most distant supernovae have shown that the expansion of space is currently accelerating (H is increasing), since around half of its life
- This can be accommodated in the equations by adding a term of pure energy of the space (not linked to any source) that has negative pressure; called "Dark Energy"
 - Analogy: space is like a spring that releases its potential energy into its expansion
- We don't know what it is, yet
- Very recent research line; attend C.Ringeval's course and the specialistic cosmology seminars to know more!

Energy composition of the Universe

